



STScI | SPACE TELESCOPE
SCIENCE INSTITUTE

EXPANDING THE FRONTIERS OF SPACE ASTRONOMY

Probing the high- z TDE population with WFIRST and JWST

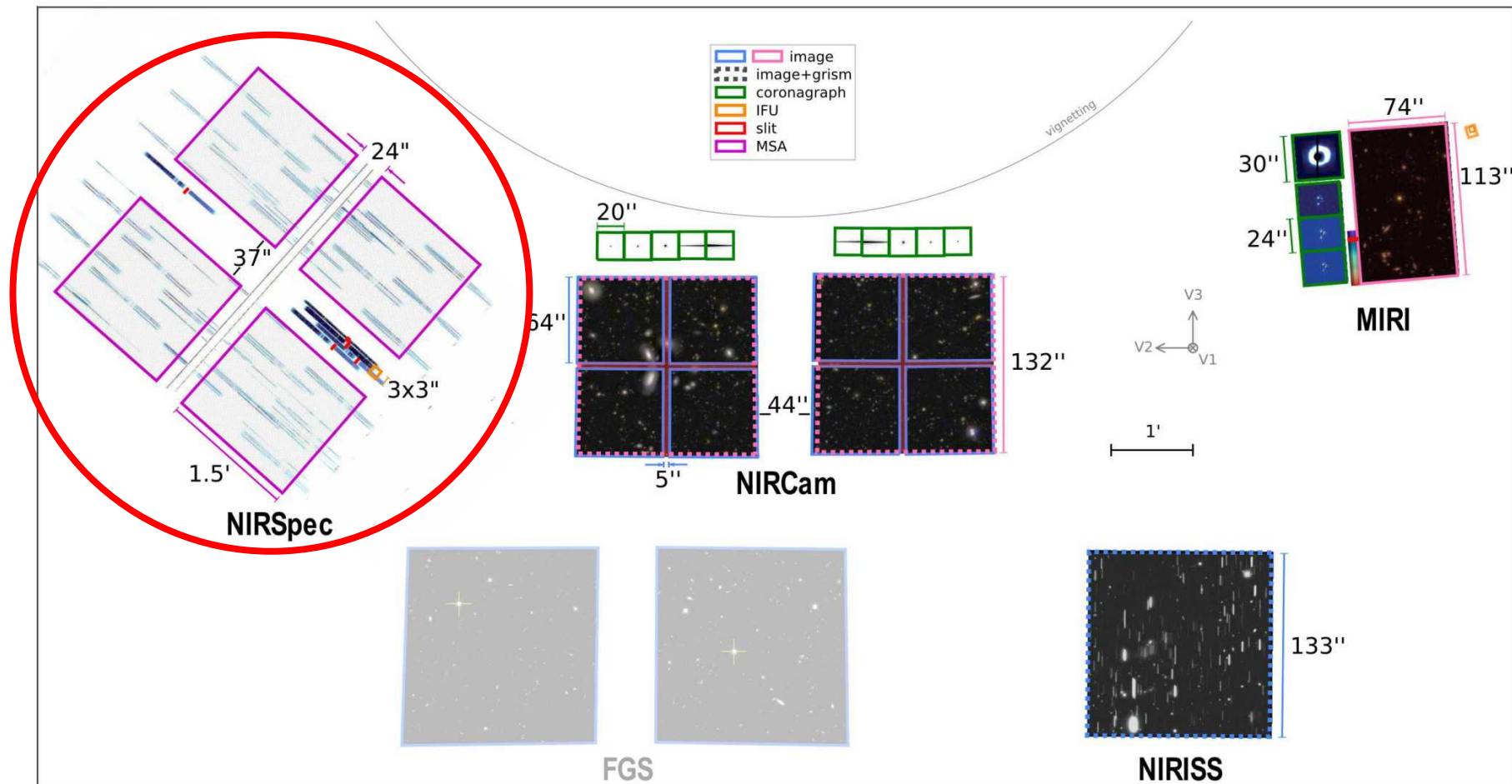
A. Rest, A. Canipe, G. Narayan (STScI),

B. R. Hounsell (UCSC), D. Scolnic (KICP), R. Chornock (UO)

Unveiling the Physics Behind Extreme AGN Variability conference, UVI July 14th 2017

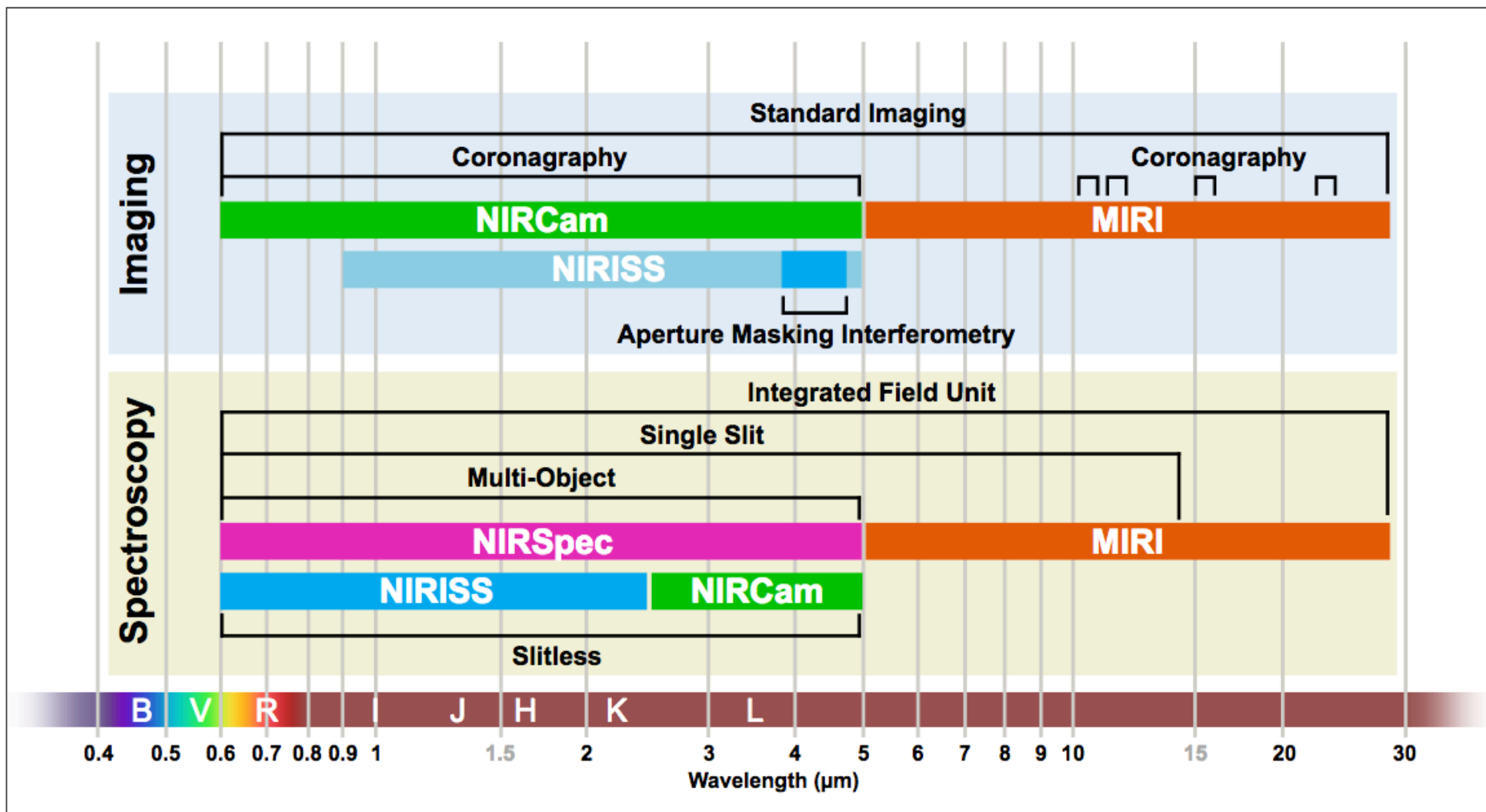


JWST Instruments



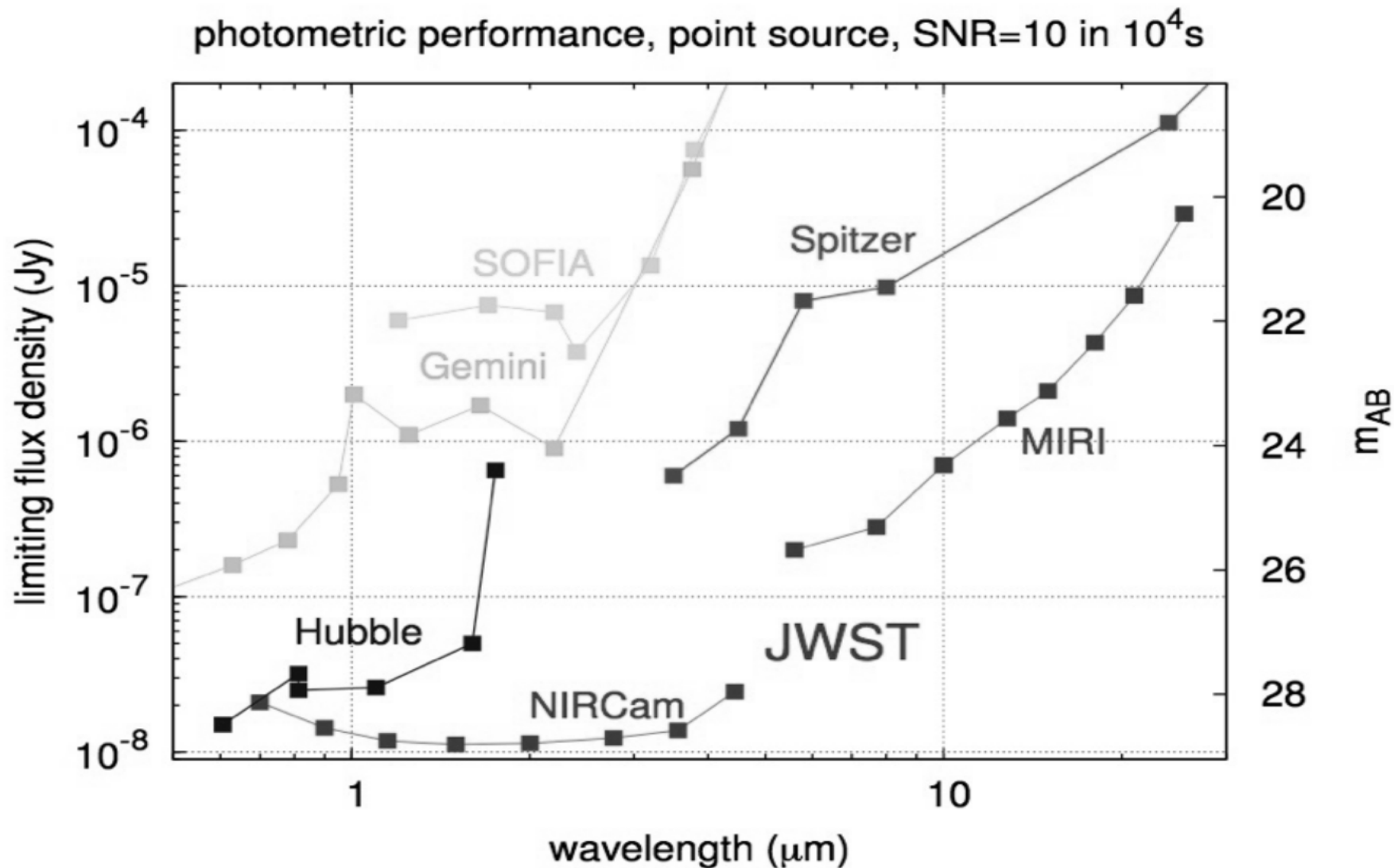


JWST: wavelength coverage



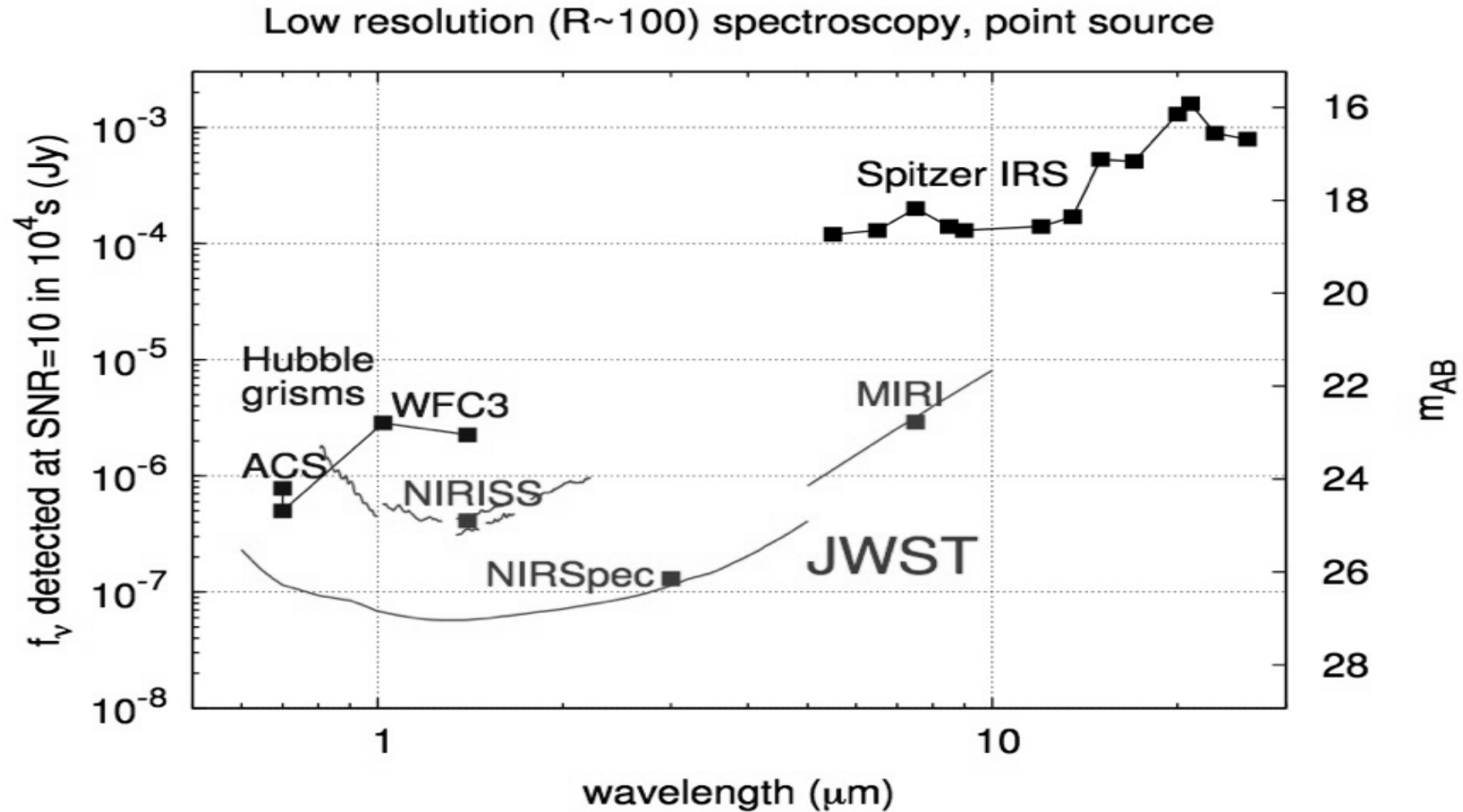


JWST: Photometric Sensitivity





JWST: Spectroscopic Sensitivity



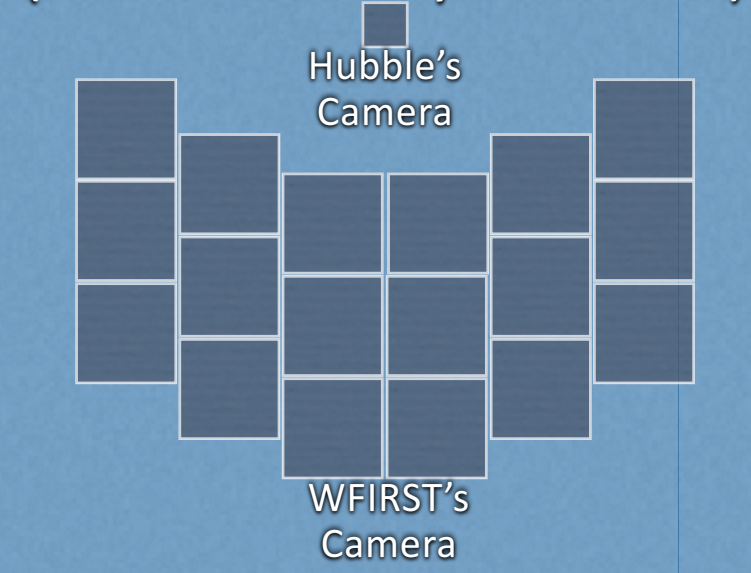


WFIRST summary

WFIRST Imaging Capabilities

Telescope Aperture (2.4 meter)	Field of View (45'x23'; 0.28 sq deg)		Pixel Scale (0.11 arcsec)		Wavelength Range (0.5-2.0 μm)		
Filters	R062	Z087	Y106	J129	H158	F184	W146
Wavelength (μm)	0.48-0.76	0.76-0.98	0.93-1.19	1.13-1.45	1.38-1.77	1.68-2.00	0.93-2.00
Sensitivity (5 σ AB mag in 1 hr)	28.50	28.02	27.95	27.87	27.81	27.32	28.33

100x the Hubble Field of View
(at the same sensitivity and resolution)



Possible Survey Implementations

High Latitude Survey (2000 sq deg at 27th mag in YJHF184 + spectra)
Dark Energy — Cosmic Lensing — High-z Galaxies — Galactic Halo Substructure

Deep Field Surveys (~10 deg² fields at 28-29th mag, with high cadence)
Supernova Discovery — First Light — Galaxy Evolution

Guest Observer Surveys (user specified instrument, depth, area, ...)
Broad astrophysics from Solar System exploration to cosmology

Galactic Bulge Survey (2.2 sq deg at high cadence)
Exoplanet Census — Free Floating Planets — Stellar Pops — Galactic Structure

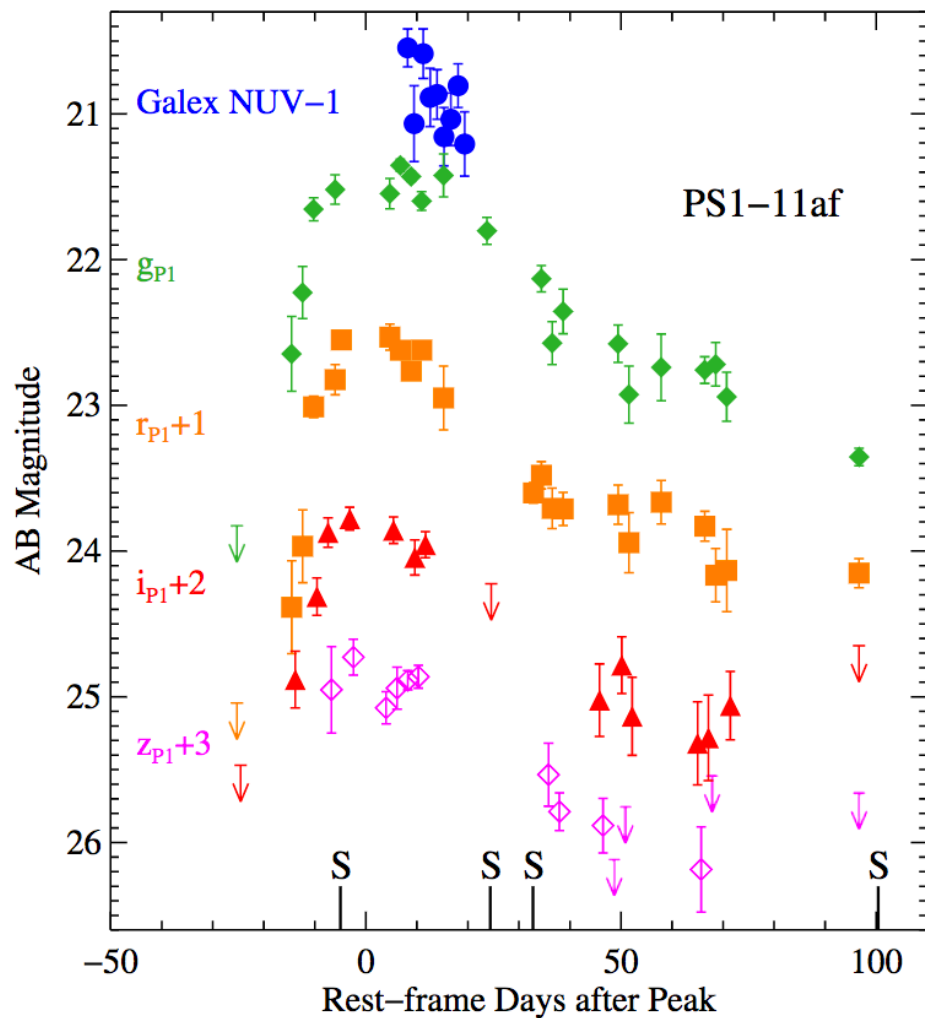
Exoplanet (+ Other Objects) Imaging Survey (10⁹ contrast ratio direct imaging)
Exoplanet Discovery and Characterization — Disks — Massive Star Atmospheres

Guest Investigator Surveys (funded archival research from survey data)
Broad astrophysics from Solar System exploration to cosmology

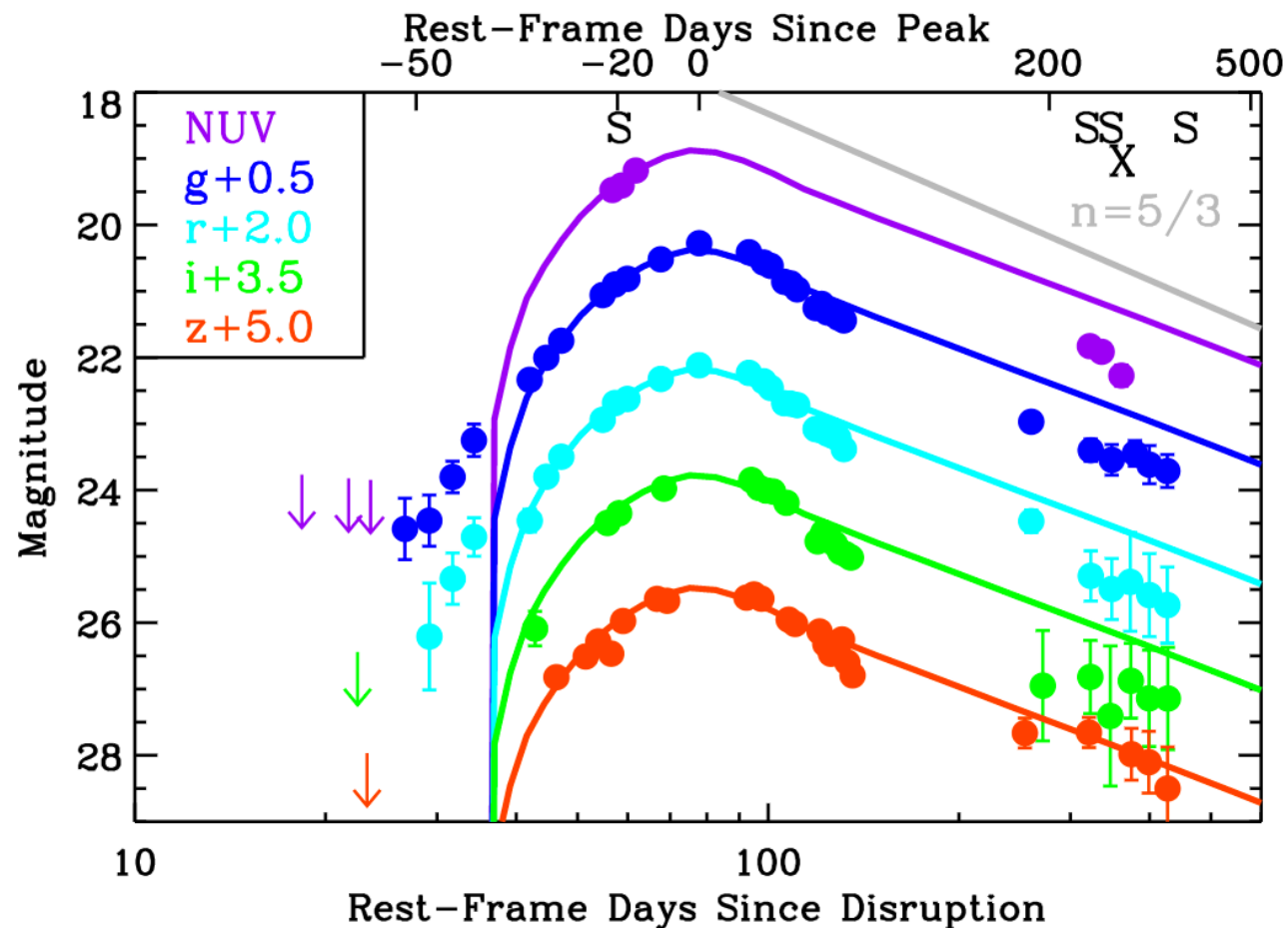


TDE Examples for Simulations: PS1-11af and PS1-10jh

Chornock+14, PS1-11af, $z=0.4$

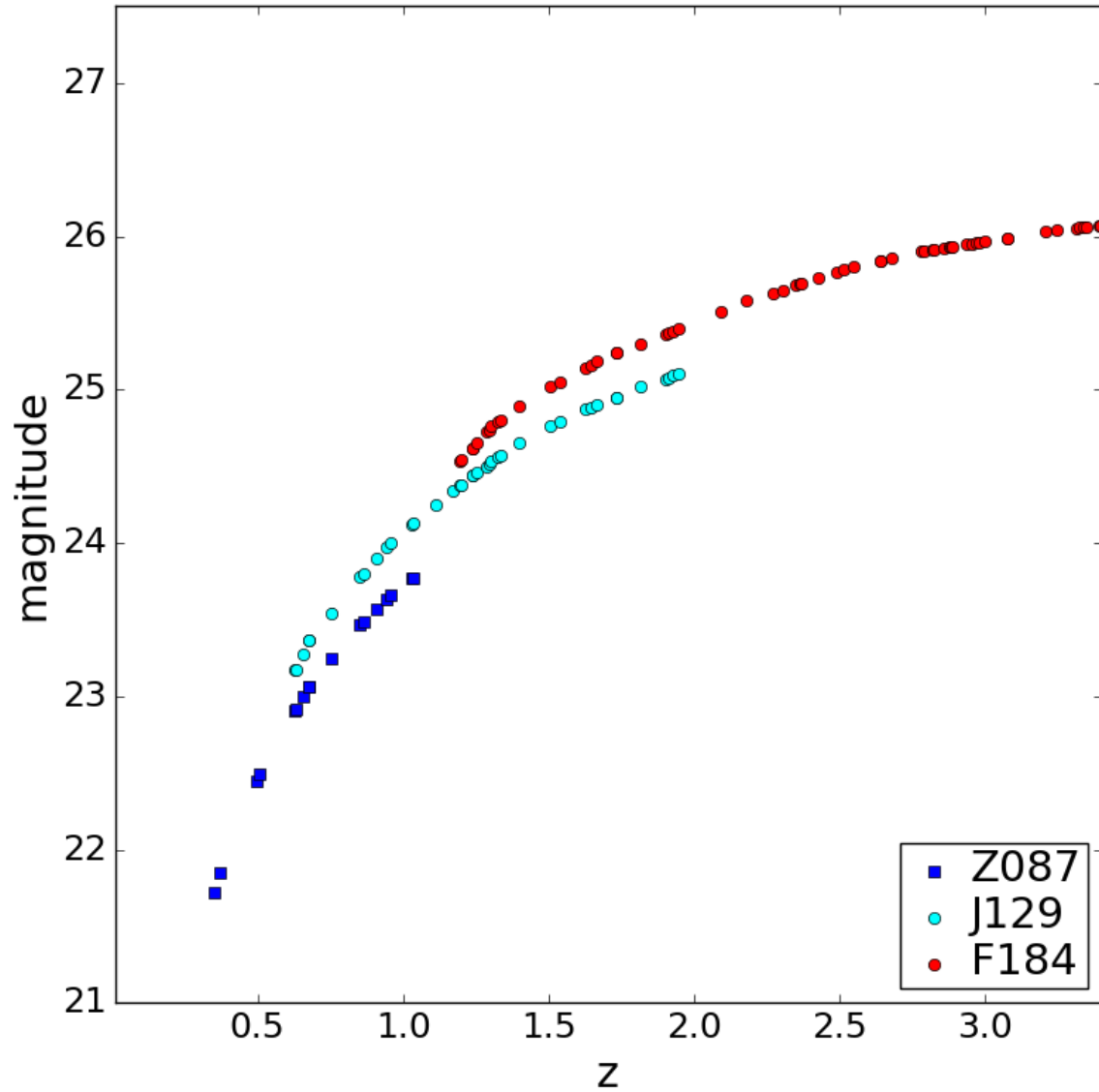


Gezari+12, PS1-10jh, $z=0.16$





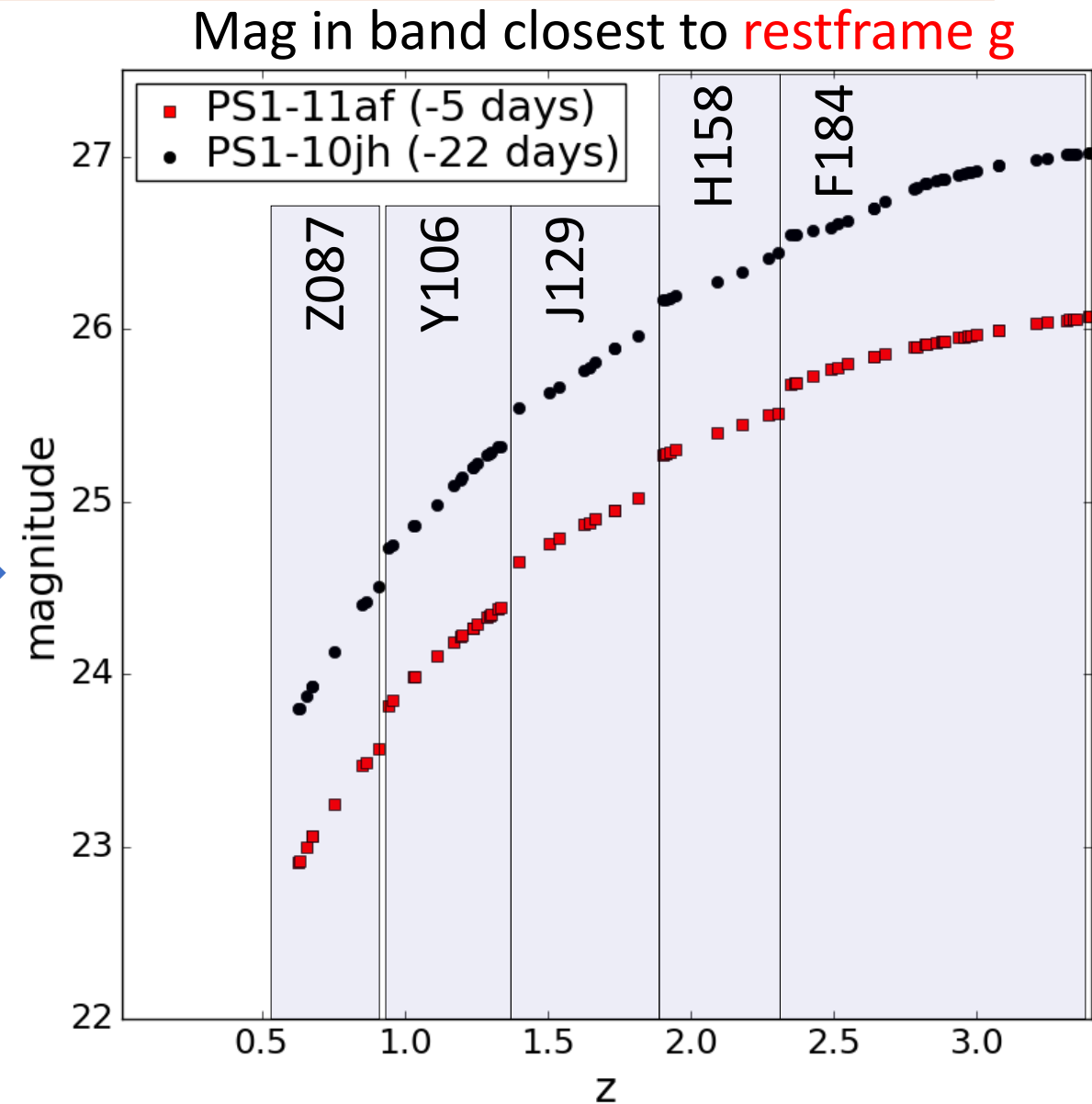
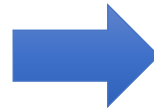
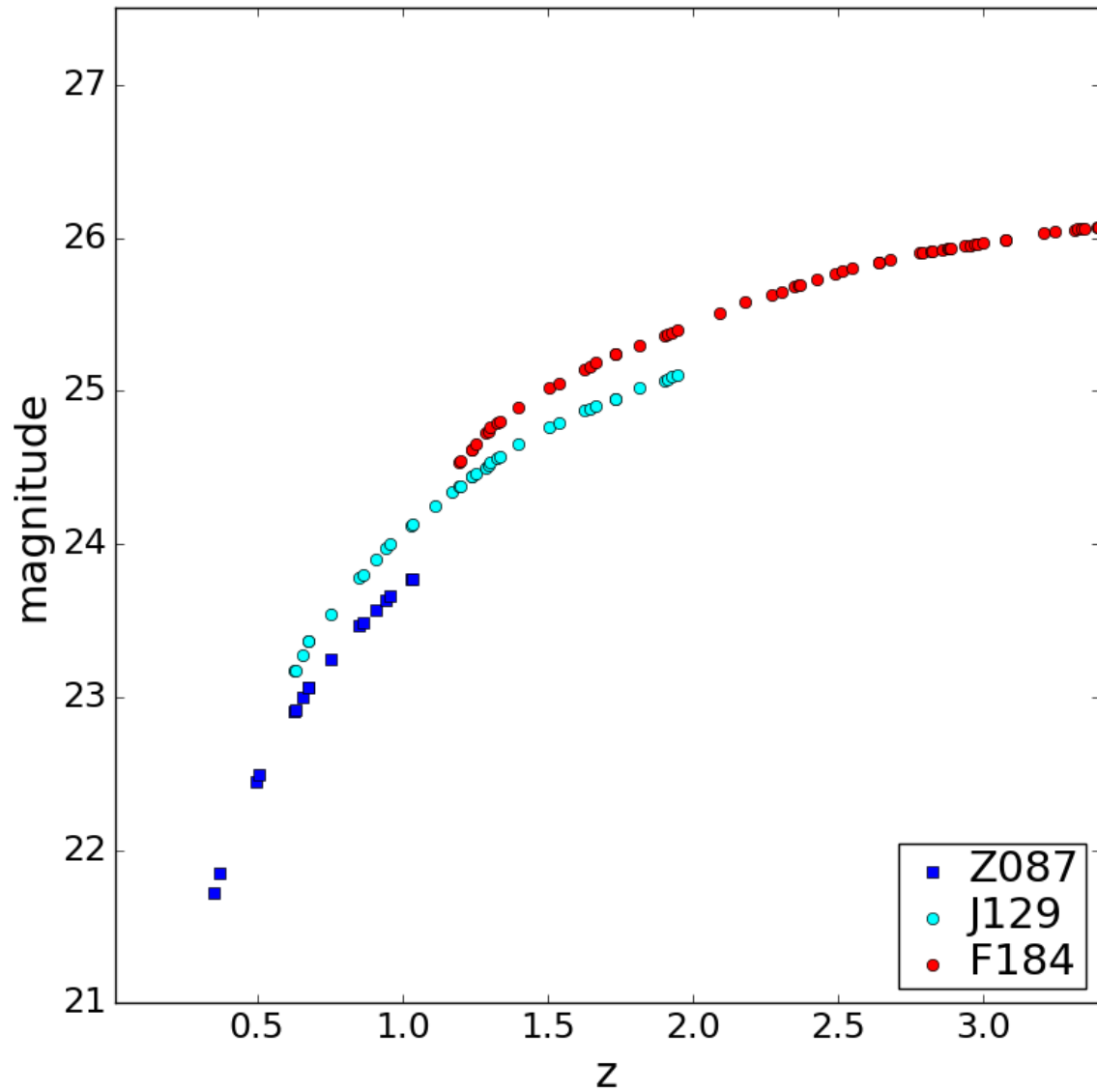
Simulating WFIRST magnitudes for PS1-11af and PS1-10jh



- SALT2
- z from 0.2 to 3.5
- Mags for given WFIRST filter



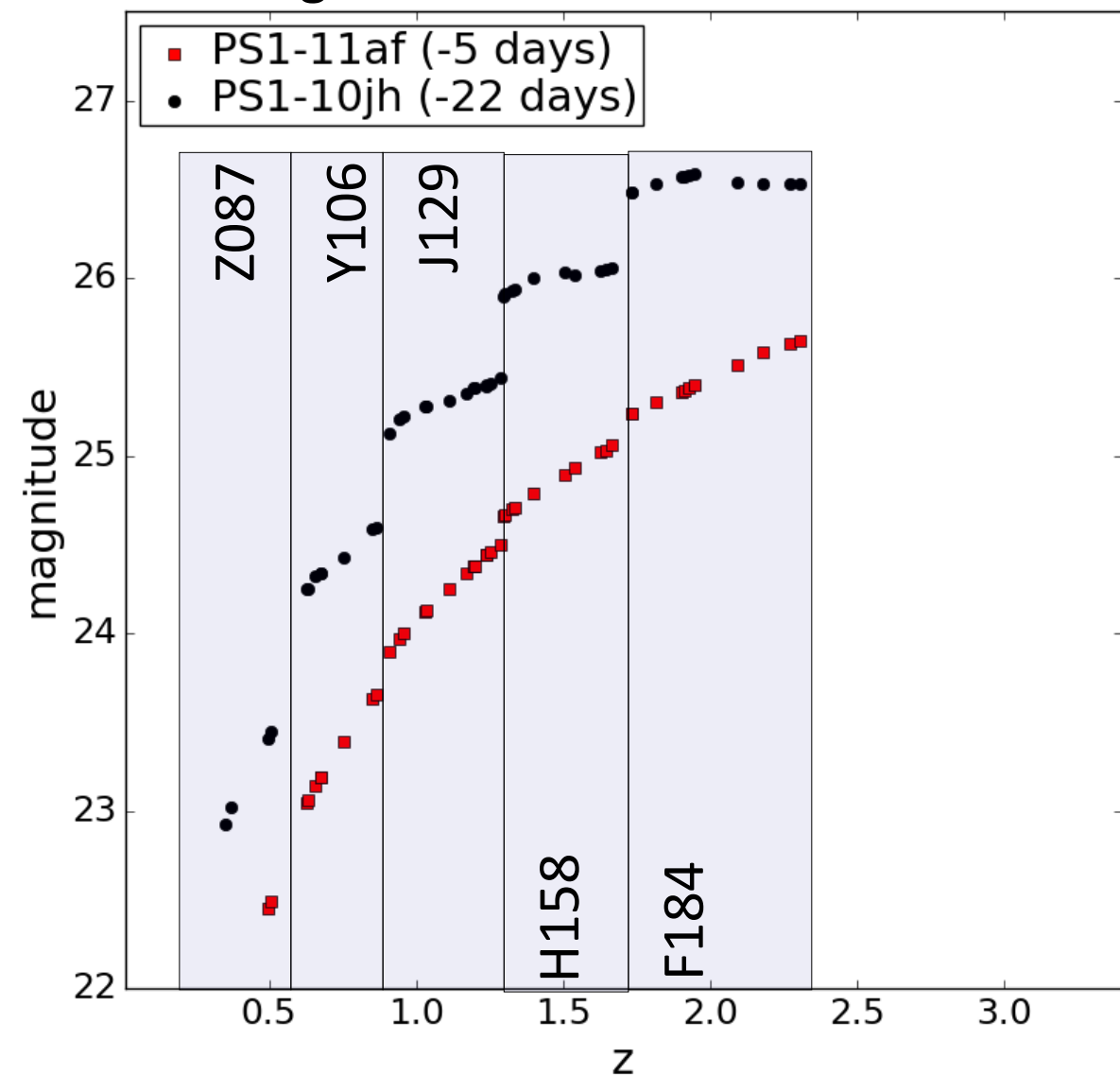
Simulating WFIRST magnitudes for PS1-11af and PS1-10jh



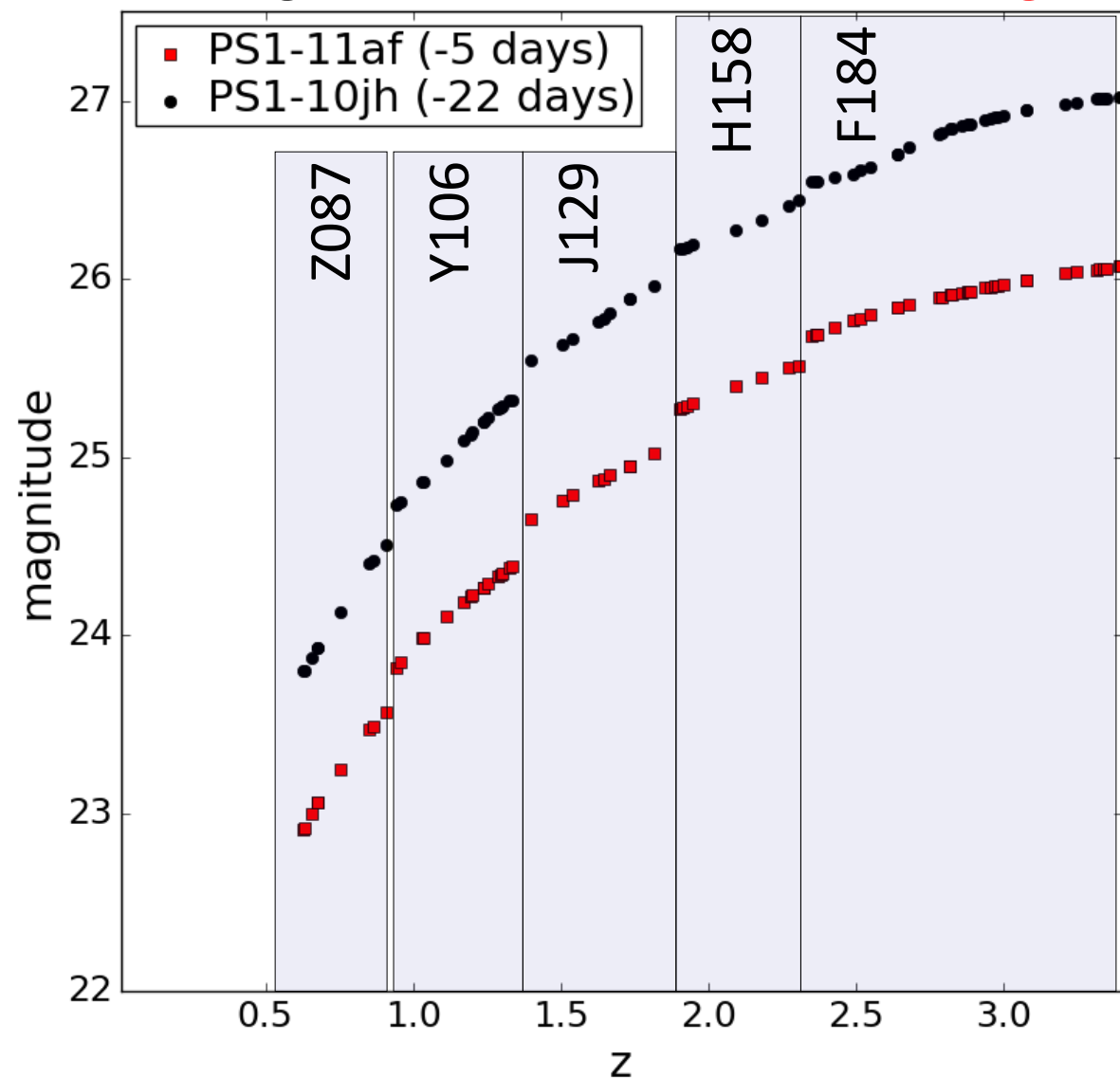


Simulating WFIRST magnitudes for PS1-11af and PS1-10jh

Mag in band closest to **restframe r**



Mag in band closest to **restframe g**





WFIRST: Survey strategies

Hounsell+17, arXiv::1702.01747

Table 2. Description of the three-tier SN survey as outlined in the SDT report.

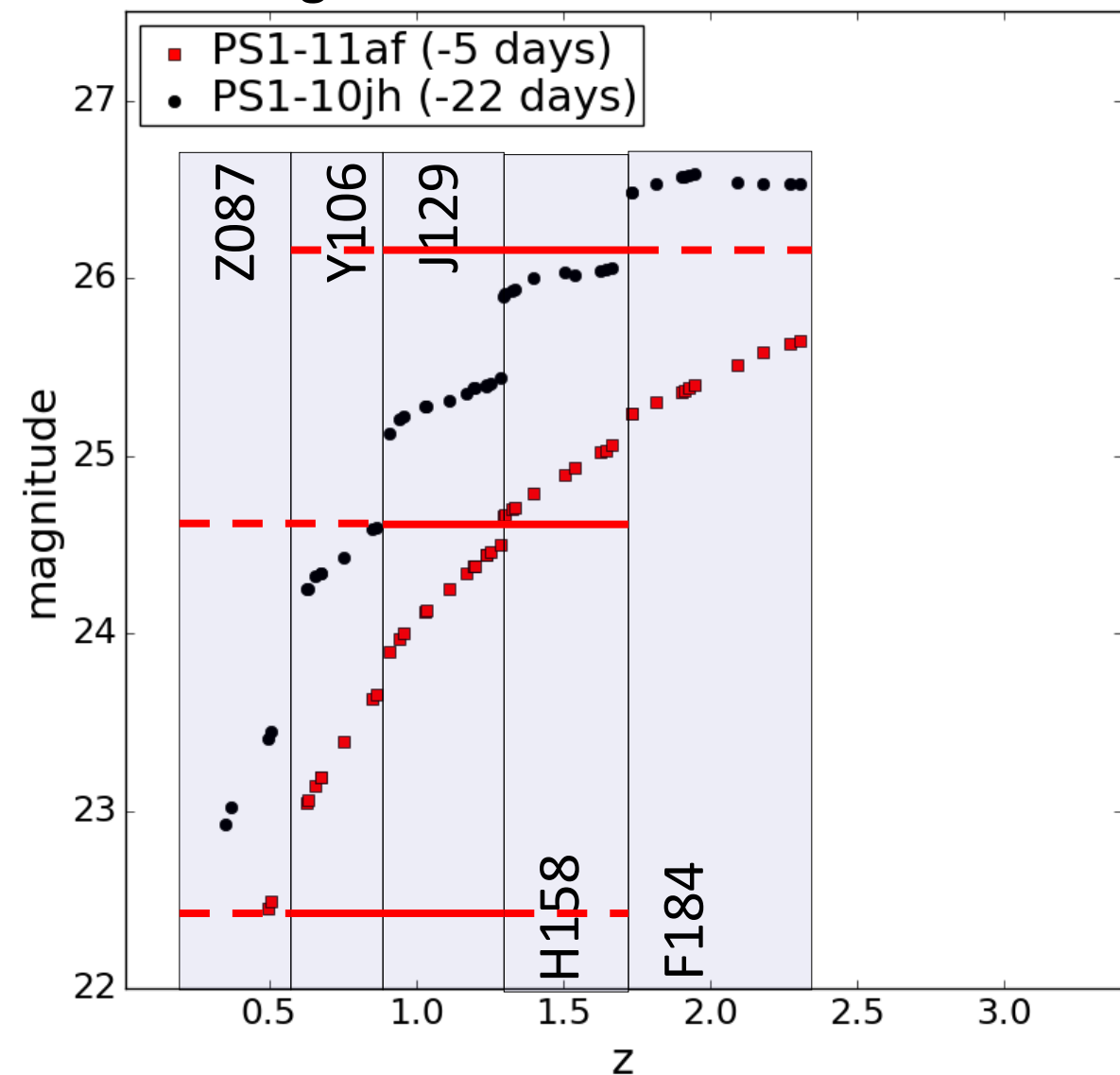
Survey	Redshift	Area	Discovery	Depth per	Total Depth
Tier	Range	(deg ²)	Filters	Exposure (mag)	(mag)
Shallow	$0.1 \leq z < 0.4$	27.44	Y, J	22.3, 22.4	25.0, 25.1
Medium	$0.4 \leq z < 0.8$	8.96	J, H	24.6, 24.5	27.3, 27.2
Deep	$0.8 \leq z \leq 1.7$	5.04	J, H	26.2, 26.1	28.9, 28.8

Name	Redshift Range			Filter Set Used			Area (deg ²)			Number of SN Ia Selected		
	Shallow	Medium	Deep	Shallow	Medium	Deep	Shallow	Medium	Deep	Shallow	Medium	Deep
SDT	0.10–0.39	0.40–0.79	0.80–1.70	IFC-S, YJ	IFC-S, JH	IFC-S, JH	27.44	8.96	5.04	12	364	1204
SDT*	0.10–0.39	0.40–0.79	0.80–1.70	IFC-S, YJ	IFC-S, JH	IFC-S, JH	27.44	8.96	5.04	149	647	1224
SDT* Highz	...	0.10–0.79	0.80–1.70	...	IFC-S, JH	IFC-S, JH	...	22.80	5.04	...	1271	1224
Imaging:Allz	0.01–2.99	0.01–2.99	0.01–2.99	$RZYJ$	$RZYJ$	$YJHF$	48.82	19.75	8.87	557	4807	5892
Imaging:Lowz	0.01–2.99	0.01–2.99	...	YJ	JH	...	142.30	66.91	...	0	1797	...
Imaging:Lowz*	0.01–2.99	0.01–2.99	...	$RZYJ$	$RZYJ$	...	73.57	32.24	...	822	8117	...
Imaging:Lowz+	0.01–2.99	0.01–2.99	...	$RZYJHF$	$RZYJHF$	...	50.66	20.68	...	588	5167	...
Imaging:Lowz-Blue	0.01–2.99	0.01–2.99	...	$BVRIYJ$	$BVRIYJ$	...	50.66	20.68	...	347	4894	...
Imaging:Highz*	...	0.01–2.99	0.01–2.99	...	$RZYJ$	$YJHF$	...	32.06	13.24	...	7990	8881
Imaging:Highz+	...	0.01–2.99	0.01–2.99	...	$RZYJHF$	$RZYJHF$	...	20.50	9.14	...	5211	6289

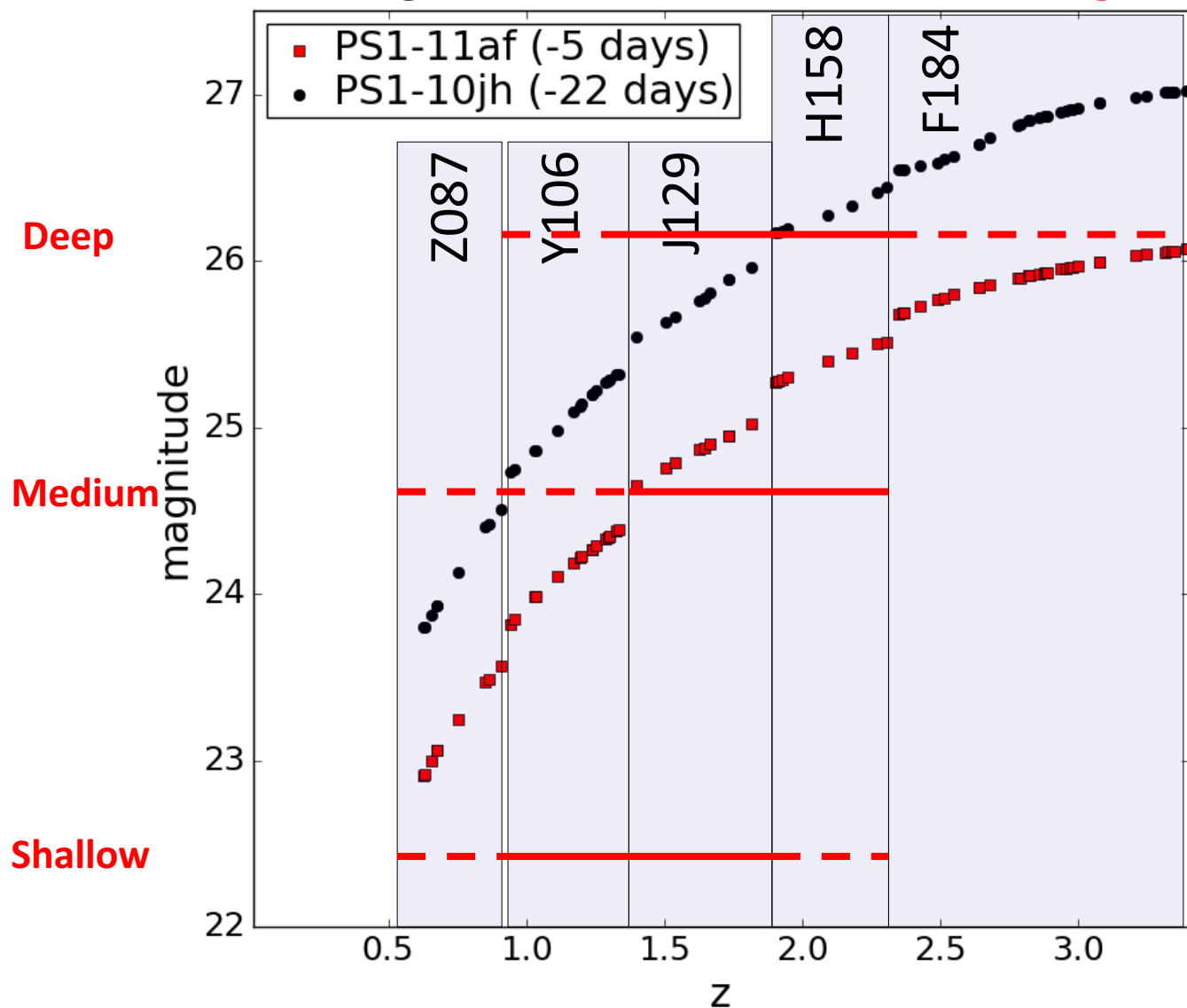


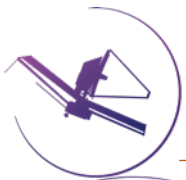
TDEs with WFIRST: survey depths

Mag in band closest to **restframe r**



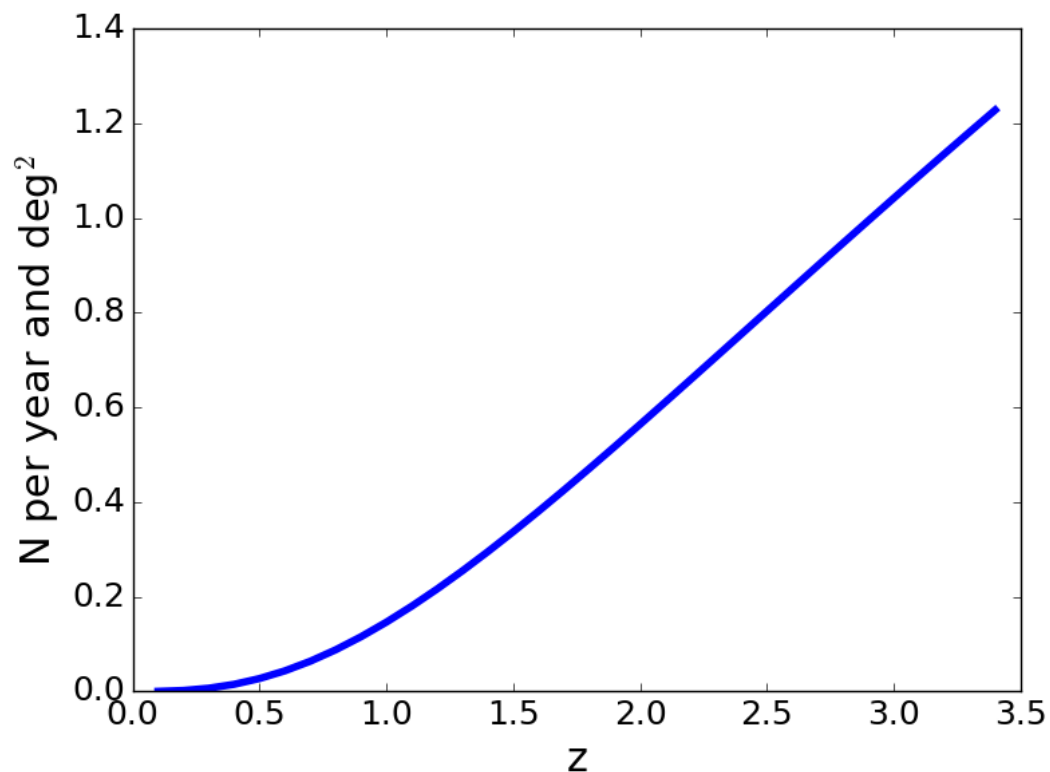
Mag in band closest to **restframe g**





TDEs with WFIRST: Rates (Assuming 4×10^{-8} TDEs $\text{Mpc}^{-3} \text{yr}^{-1}$ (van Velzen+14))

Survey name	Survey Area Range (deg ²)	z max	N TDE (yr ⁻¹)
Shallow	25 - 50	<0.5	?
Medium	10 - 20	<1.5	3 - 6
Deep	5 - 10	<3.0	2 - 8

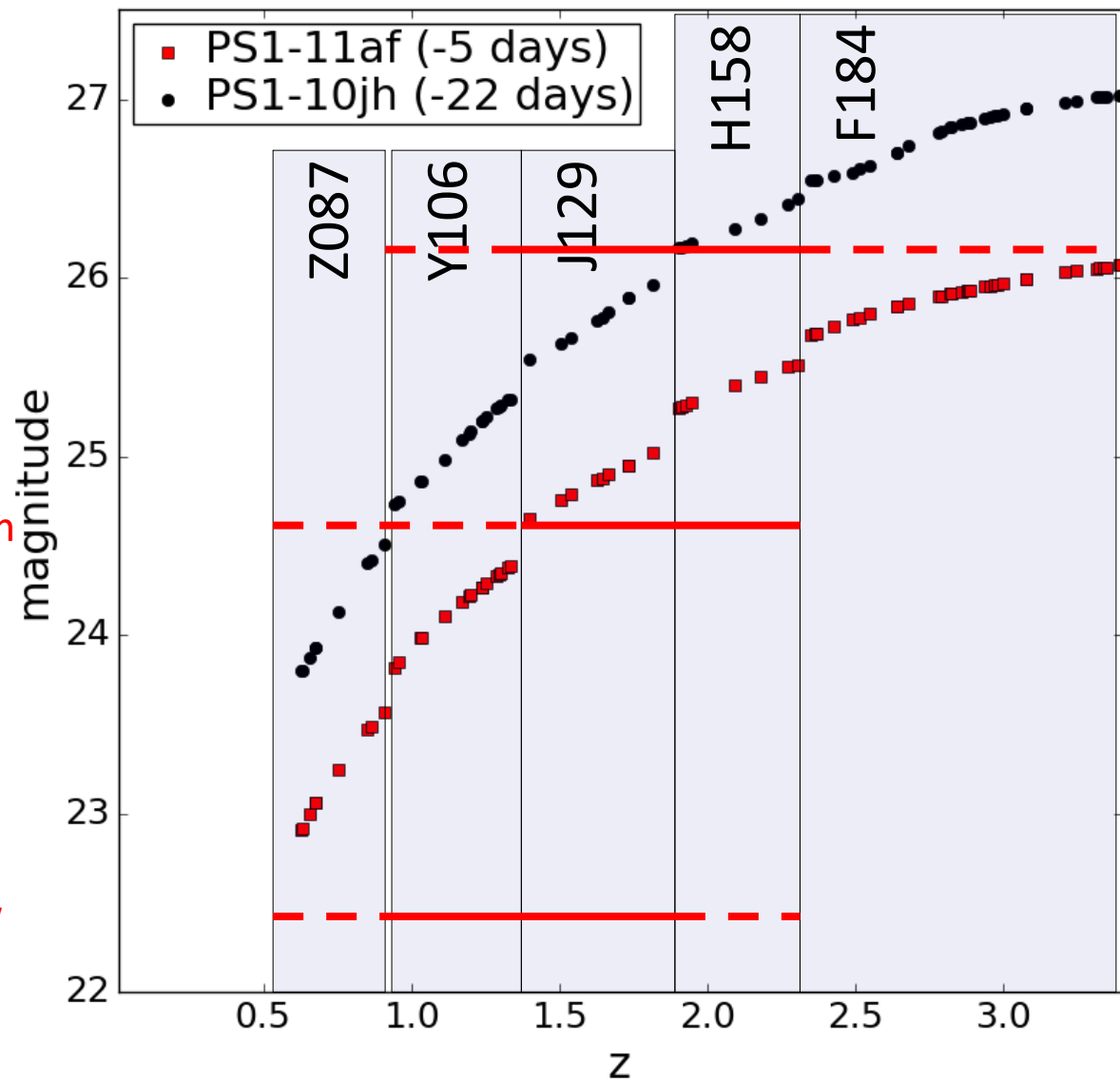


Deep

Medium

Shallow

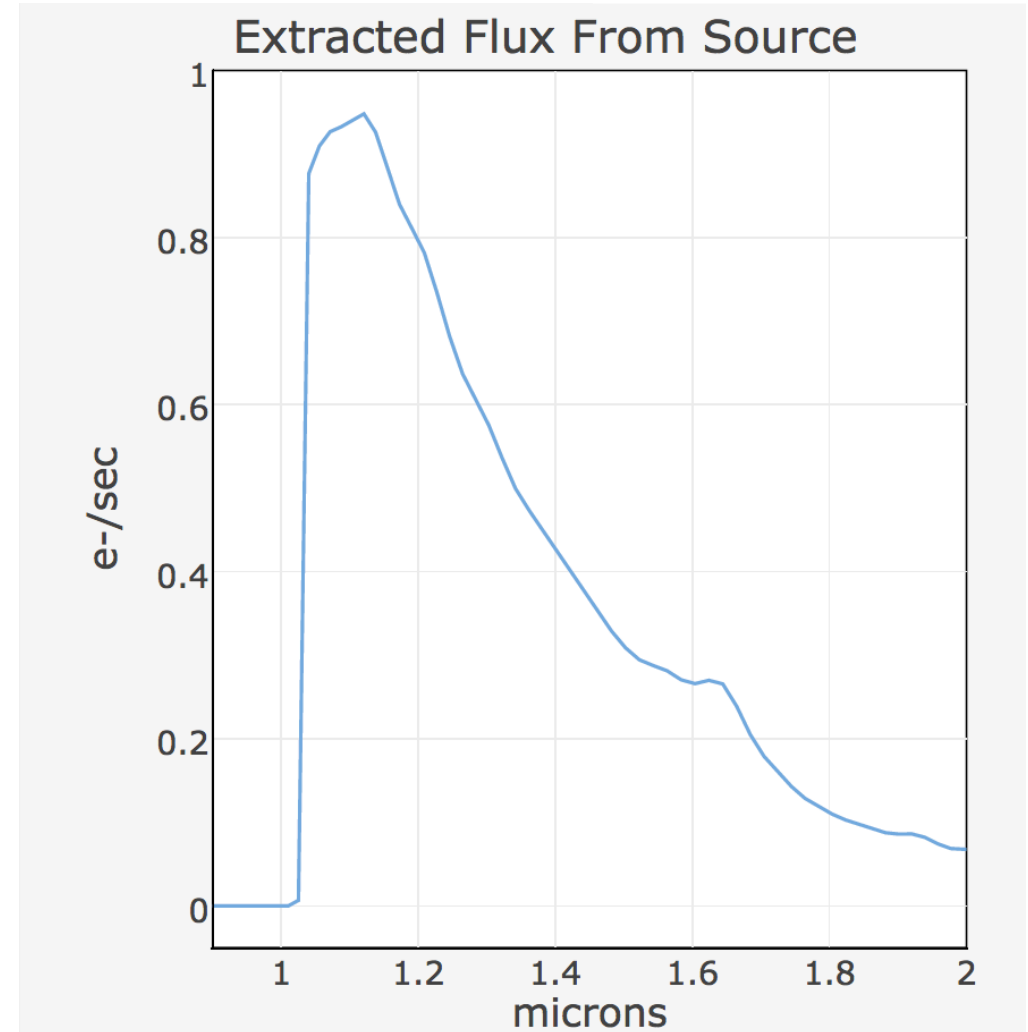
Mag in band closest to **restframe g**





TDEs at high redshift: Spectroscopy with JWST NIRSpec

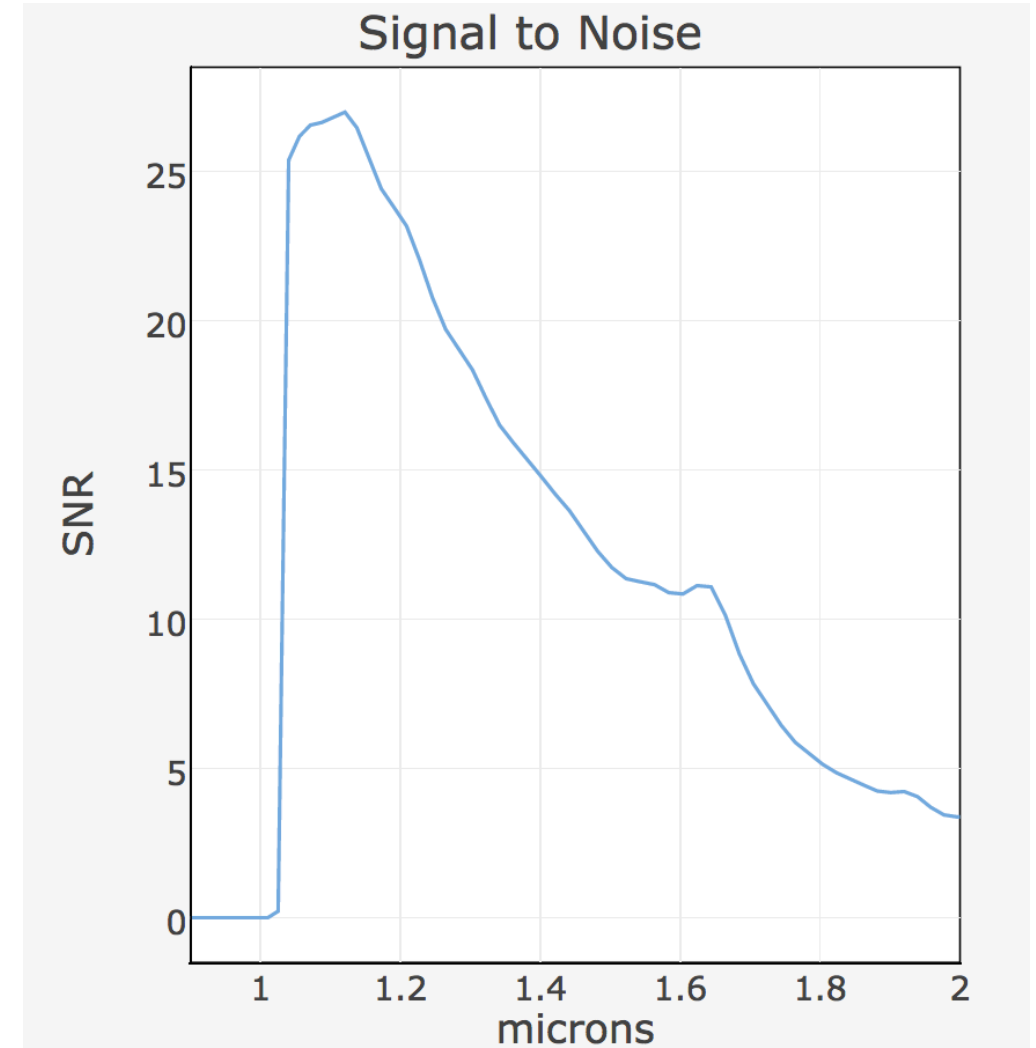
- JWST Exposure time calculator
- PS1-10jh, phase=-22 spectrum
- Normalization
 - $z=2.5$
 - F150W=25.7
 - 6000 seconds exposure time
- PRISM/CLEAR
- S200 A1 (0.2" x 3.3")
- 0.6 - 5.3 microns
- $R=100$
- CAVEAT: Does not include host galaxy!





TDEs at high redshift: Spectroscopy with JWST NIRSpec

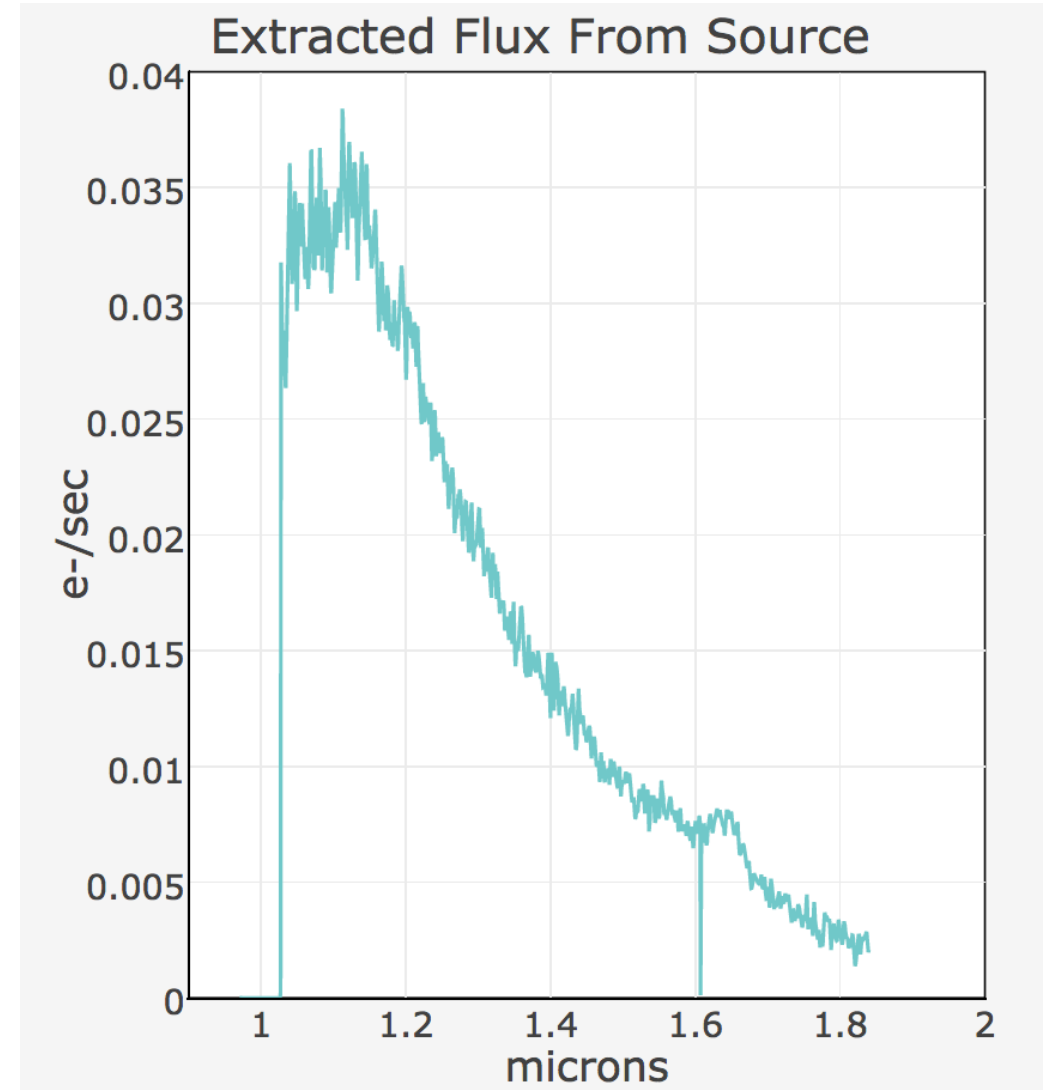
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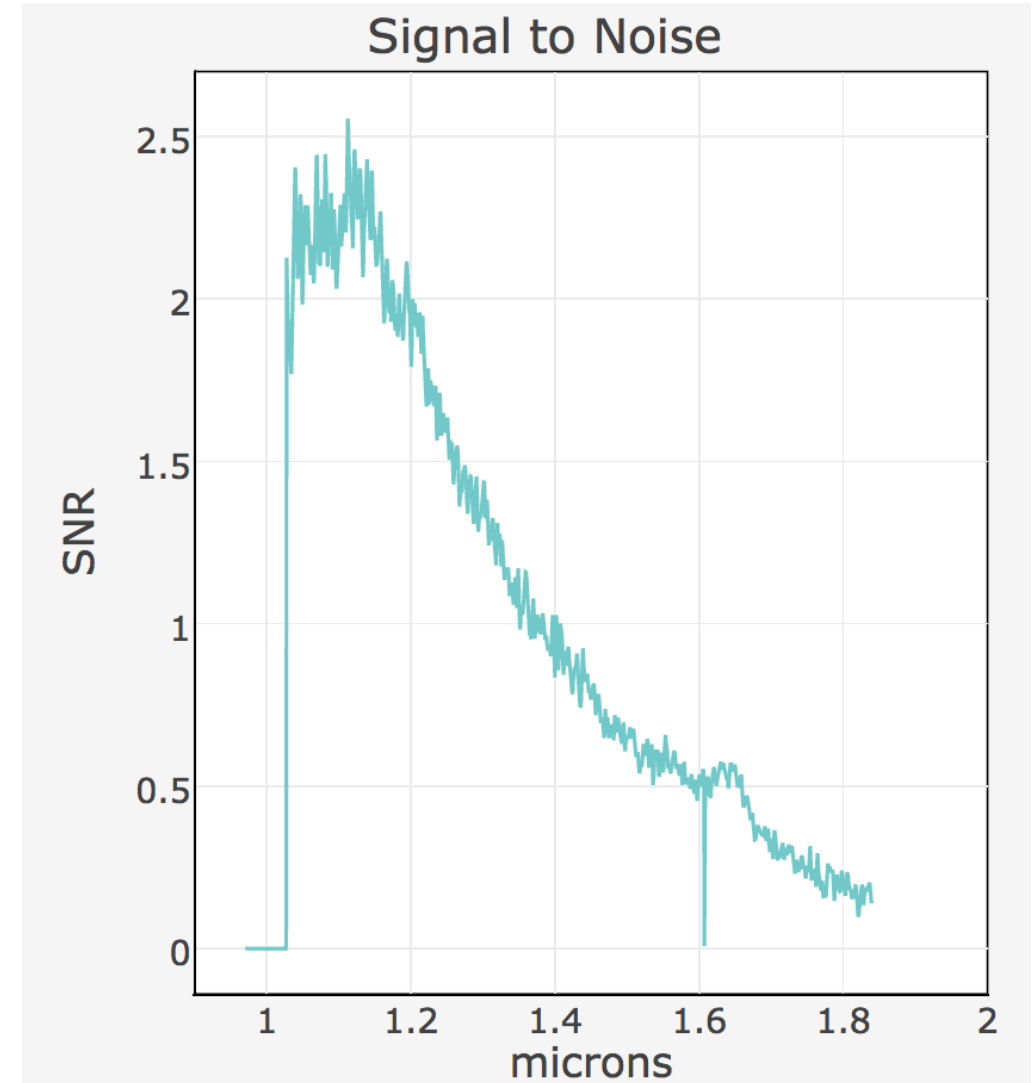
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Summary

- Low-z TDE sample will grow from 10s to hundreds (ZTF) and thousands (LSST)
- Simulation: Use PS1-10jh and PS1-11af as examples
- High-z TDE sample with WFIRST:
 - Restframe g observable up $z=3.5$ with F184
 - SN survey: 5-10 TDE within $z=3.5$ per year (Efficiency? Constant rate with redshift?)
- JWST:
 - $R=100$ spectroscopy easily achievable for all discovered TDEs
 - $R=1000$: also possible!
- Work in progress, paper planned for late summer:
 - TDEs
 - SN Ia
 - SN IIP
 - SN Ib/c
 - SLSN